## **Data Interpretation and projects**





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## **OUTLINE**

- Near midnight TEC enhancements [Valladares and Chau, Radio Sci., 2012].
- Detection of Gravity waves (TIDs). Large-scale GWs. Middle-scale GWs measured using GPS interferometry and cross-correlation techniques [Valladares and Hei, IJG, 2012].
- TEC depletions and scintillations [Seemala and Valladares, Radio Sci., 2011].
- Tropical TEC enhancements [Valladares and Eccles, 2013].



#### 24-hr TEC values for 16 consecutive days from Huancayo



## Latitudinal Variability of TEC due to anomaly



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#### 24-Hour plots for several stations during the near-midnight TEC enhancement



#### Jicamarca density observations for the same day (March 11-12, 2011)



#### ΓEC variability over South America measured at 01, 03, 04, and 05 UT



The combination of upward, daytime ExB drift velocity perpendicular to B and downward diffusion parallel to B by gravity and pressure gradient forces create crests in ionization at +/- 15 to 20 degrees magnetic latitude known as the equatorial anomaly. If the daytime, ExB drift velocities are significantly lower or are absent, then the crests in ionization are significantly closer to the magnetic equator or are absent



Provided by D. Anderson

Low Latitude Transport Mechanisms

## **TEC values along 3 field lines**



#### **How TEC values are filled in: TEC for October 16, 2008**



Geographic Latitude

## **Daytime double peak (Kp = 7+)**





#### **TEC values during magnetically disturbed conditions**



**Atmospheric Gravity Wave Origin and Dynamics** 

GW are produced by any disturbance in the atmosphere on a time scale of a few minutes to several hours. The restoring force is gravity.

Tropospheric disturbances such as: weather fronts, depressions, jet streams, severe storms, wind blowing over topographical features. – And, nuclear explosions.

Intense Joule heating deposited at high latitudes during Super magnetic storms (data from the Oct 2003 super storm to be presented).



#### ALTAIR Zonal Scan [Hysell et al., Ann. Geophys. 2006]

Sat Aug 7 08:40:04 2004



## **TEC wave Perturbations associated with TIDs**



#### **GW/TIDs Observations using high-resolution VIPIR ionograms**

High Resolution E-region meas. around sunrise

**E-region altitude** 





Measurements of large-scale gravity waves during the super storm of Oct, 2003 (Halloween storm)

## Large-scale TEC variations seen in South America

#### on October 29,2003

TEC variations with time scales of many tens of minutes produced by the passage of Atmospheric Gravity Waves (AGW)



SITEC, sudden increase in TEC (associated with solar flares)

#### **Large-scale TEC perturbations at Lat=34°N, Long=80°W.**



## **TEC perturbation in North America (October 29, 2003)**



Dedicated campaign to detect atmospheric gravity waves using 3 GPS receivers

## **GPS Radio-interferometry**

Radio interferometers are arrays of two or more ground receivers that use phase differences measured at various stations to determine TID velocity, propagation azimuth, amplitude, and period.





For geo-stationary spatial and time derivatives of the phases give phase interference pattern. GPS radiointerferometry uses the spatial and temporal derivatives of TEC.

Afraimovich et al., 1998; Afraimovich et al., 2002

#### **Location of 3 GPS receivers near Huancayo**



## **Formulas**

It is assumed that the TEC perturbation (gravity wave packet) follows the formula:

 $\Phi(\text{TEC}) = \delta \sin(\Omega t - KxX - KyY + \varphi)$ 

The Angle of arrival  $\alpha(t)$ , in the horizontal plane, is given by:  $\alpha(t) = \arctan \left( \frac{G_x(t)}{G_y(t)} \right) = \arctan \left( \frac{\phi'_x(t)}{\phi'_y(t)} \right)$ (5)

Phase velocity  $\mu(t)$  in GPS satellite frame of reference:  $u_x(t) = \phi'_t(t)/\phi'_x(t) = u(t)/\sin\alpha(t)$  $u_y(t) = \phi'_t(t)/\phi'_y(t) = u(t)/\cos\alpha(t)$  $u(t) = |u_x(t)u_y(t)|(u_x^2(t) + u_y^2(t))^{-1/2}$  (8)

## **TEC values measured at Huancayo on July 20, 2008**



#### **TEC perturbations for PRNs 22 and 32 recorded on July 20, 2008**



## **TIDs observed on July 20, 2008**



#### **X-Correlation Functions for HYO – Cuzco and HYO - Piura**



Map of TEC perturbations (associated with TIDs) in South America



In 1 hour this patch of TIDs moved further north

## GW phase velocities measured on July 20, 2008 during a 3-week campaign



Phase velocity calculated at Huancayo using 3 closedspaced GPS receivers. Similar velocities were measure by Valladares and Chau, [2012] using GPS placed hundreds of km apart.

## TEC perturbations after detrending daily variability

Satellite passes for several days are shown in different colors to emphasize dayto-day variations.



## **Mountain Waves near Concepcion, Chile**



**Slide Provided by Wes Swartz** 

## Satellite 11 Path and Andes Topography

The distance from the satellite to the Andes ridge line varies quite a bit from -35° to -30° latitude.

Apparently, the varying distance has little effect on undulations.

The altitude of the ridge line has the most effect.



#### Satellite 08 Path and TEC perturbations at Iquitos



#### **TEC values for July 12, 2011 (quiet conditions)**



#### TEC values for June 7, 2011 between 12 and 23 LT.



#### **TEC values for December 12, 2011**





#### Wind system associated with tropical TEC enhancement

Operates during the June Solstice

Both the westward & Northsouth wind move plasma up the field lines.



# TEC results of the numerical model LLIONS using 90 planes along parallel field lines. One every degree between 120° W and 30° W.





## Conclusions

- Networks of GPS receivers are the best tool to make new discoveries on space weather.
- TEC values are very variable depending on solar and magnetic conditions.
- LISN is a distributed observatory to study some aspect of space weather (plasma bubbles, ESF). It provides regional coverage of the day-to-day variability of the ionosphere over South America.

- 1) Fit a 4th order polynomial to every 3 hours of TEC data.
- 2) Find the difference between TEC and fitted TEC : dTEC[1].
- 3) Determine cases when dTEC[1] is below a threshold level, say 0.7 TECu.
- 4) Make a second fit to TEC, but avoid periods when dTEC[1] is below the threshold value.
- 5) Find the difference between TEC and the second fit : dTEC[2].
- 6) Search for the beginning and end of depletions based on the derivatives of original TEC.
- 7) A low pass filter is also used to eliminate periods < 3 min.

#### **TEC depletions seen on 5 stations aligned east-west**



## **Temporal Evolution of a Gravity Wave Packet**

Due to wave dispersion, a gravity wave packet spreads out to a large volume in thermosphere. Spatial extent of wave packet at z=200-250 km is ~500-1000 km



Convective plume envelope 20 km x 20 km x 10 km

Gravity Waves move upwards and away from the source region, carrying energy and momentum flux.